

The ϕ meson at finite density, revisited

Workshop on J-PARC hadron physics in 2014

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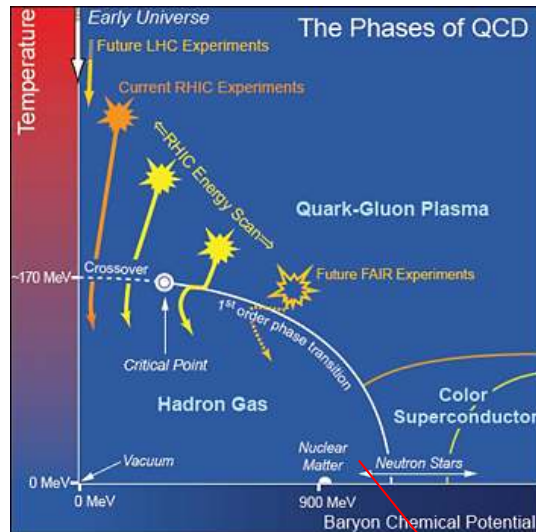
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Introduction: Vector mesons at finite density

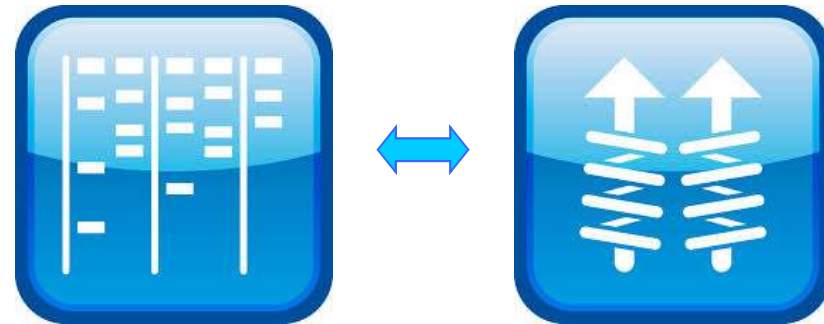
Basic Motivation:

Understanding the behavior of matter under extreme conditions



$\rho, \phi, \omega, D \dots$

Understanding the origin of mass and its relation to chiral symmetry of QCD



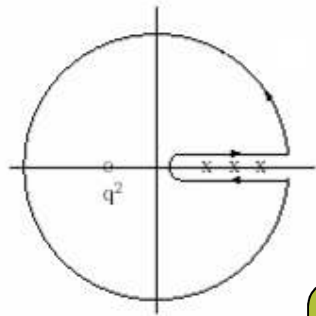
- Vector mesons: clean probe for experiment
- To be investigated at J-PARC
- Firm theoretical understanding is necessary for interpreting the experimental results!

QCD sum rules

M.A. Shifman, A.I. Vainshtein and V.I. Zakharov,
Nucl. Phys. B147, 385 (1979); B147, 448 (1979).

In this method the properties of the two point correlation function is fully exploited:

$$\Pi(q^2) = i \int d^4x e^{iqx} \langle 0 | T \{ \chi(x) \bar{\chi}(0) \} | 0 \rangle$$



$$\rightarrow \Pi(q^2) = \frac{1}{\pi} \int_0^\infty ds \frac{\text{Im} \Pi(s)}{s - q^2 - i\epsilon}$$

is calculated
“perturbatively”,
using OPE

spectral function
of the operator χ

After the Borel transformation:

$$G_{OPE}(M) = \frac{1}{\pi} \int_0^\infty ds \frac{1}{M^2} e^{-\frac{s}{M^2}} \text{Im} \Pi(s)$$

More on the OPE in matter

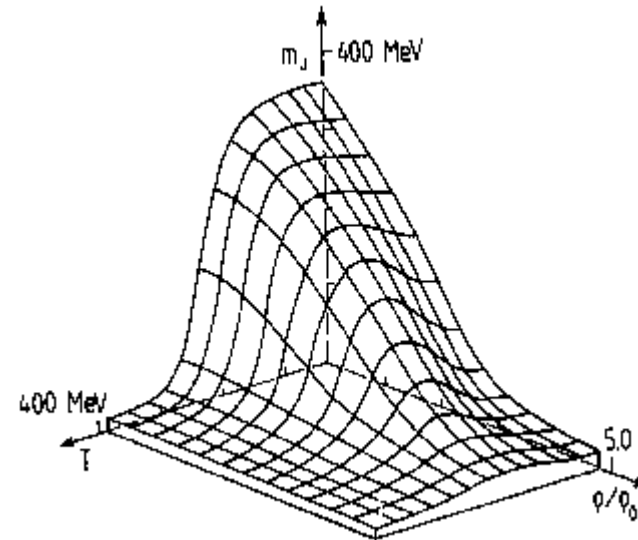
perturbative Wilson coefficients

non-perturbative condensates

$$i \int d^4x e^{iqx} \langle 0 | T \{ \chi(x) \bar{\chi}(0) \} | 0 \rangle = C_I(q^2) I + \sum_n C_n(q^2) \langle 0 | O_n | 0 \rangle$$

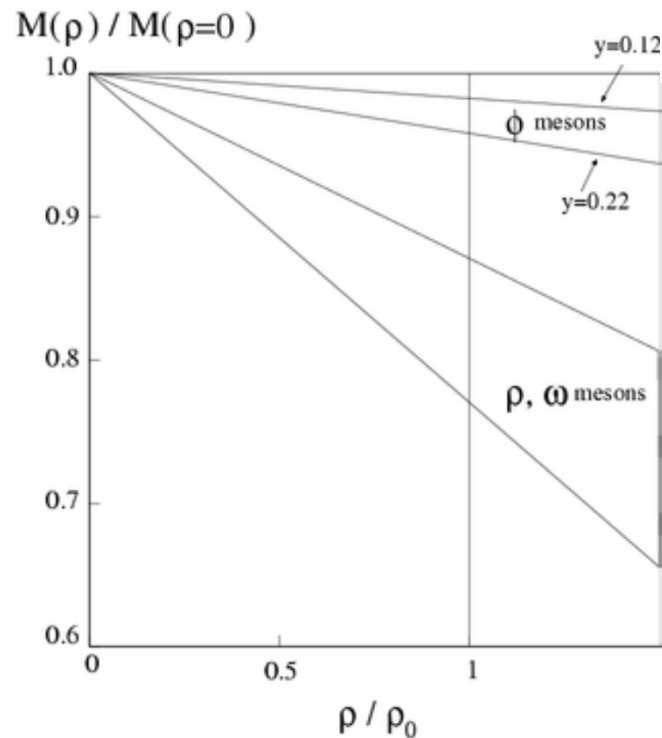
$$\begin{aligned} \langle 0 | O_n | 0 \rangle = & \langle 0 | \bar{q}q | 0 \rangle, \\ & \langle 0 | G_{\mu\nu}^a G^{a\mu\nu} | 0 \rangle, \\ & \langle 0 | \bar{q} \sigma_{\mu\nu} \frac{\lambda^a}{2} G^{a\mu\nu} q | 0 \rangle, \\ & \langle 0 | \bar{q}q\bar{q}q | 0 \rangle, \dots \end{aligned}$$

Change in hot or
dense matter!



Important early study

T. Hatsuda and S.H. Lee, Phys. Rev. C **46**, R34 (1992).



Vector meson masses mainly drop due to changes of the quark condensates.

The most important condensates are:

$$\langle \bar{q}q\bar{q}q \rangle_\rho \quad \text{for} \quad \rho, \omega$$

$$m_s \langle \bar{s}s \rangle_\rho \quad \text{for} \quad \phi$$

Important assumption:

$$\langle \bar{q}q\bar{q}q \rangle_\rho = \langle \bar{q}q \rangle_\rho^2, \quad \langle \bar{s}s\bar{s}s \rangle_\rho = \langle \bar{s}s \rangle_\rho^2$$

} Might be wrong!

Structure of QCD sum rules for the light vector mesons

$$\frac{1}{M^2} \int_0^\infty ds e^{-\frac{s}{M^2}} \rho(s) = c_0 + \frac{c_1}{M^2} + \frac{c_2}{M^4} + \frac{c_3}{M^6} + \dots$$

Vacuum

ρ, ω

$$c_0 = 1 + \frac{\alpha_s}{\pi}$$

$$c_1 = -6m_q^2 \sim 0$$

$$c_2 = \frac{\pi^2}{3} \langle \frac{\alpha_s}{\pi} G^2 \rangle + 8\pi^2 m_q \langle \bar{q}q \rangle$$

$$c_3 = -\frac{896}{81} \kappa \pi^3 \alpha_s \langle \bar{q}q \rangle^2$$

ϕ

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$$c_1 = -6m_s^2$$

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Finite density effects

Modification of condensates

$$\langle \bar{q}q \rangle_\rho = \langle \bar{q}q \rangle_0 + \langle N | \bar{q}q | N \rangle \rho = \langle \bar{q}q \rangle_0 + \frac{\sigma_{\pi N}}{2m_q} \rho$$

$$\begin{aligned} \langle \bar{s}s \rangle_\rho &= \langle \bar{s}s \rangle_0 + \langle N | \bar{s}s | N \rangle \rho = \langle \bar{s}s \rangle_0 + \frac{\sigma_{sN}}{m_s} \rho \\ &= \langle \bar{s}s \rangle_0 + y \frac{\sigma_{\pi N}}{2m_q} \frac{\langle N | \bar{s}s | N \rangle}{\langle N | \bar{q}q | N \rangle} \end{aligned}$$

$$\langle \frac{\alpha_s}{\pi} G^2 \rangle_\rho = \langle \frac{\alpha_s}{\pi} G^2 \rangle_0 - \frac{8}{9} (M_N - \sigma_{\pi N} - \sigma_{sN}) \rho$$

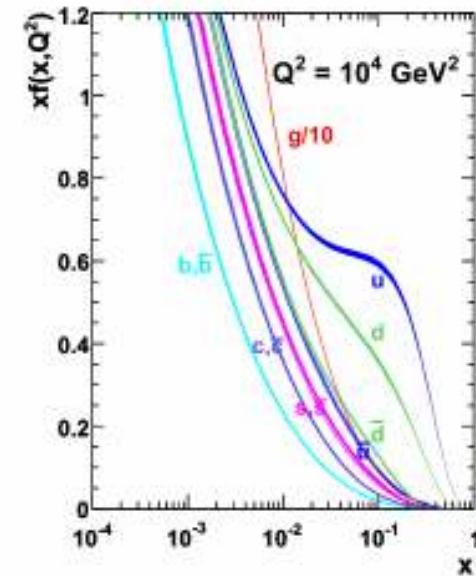
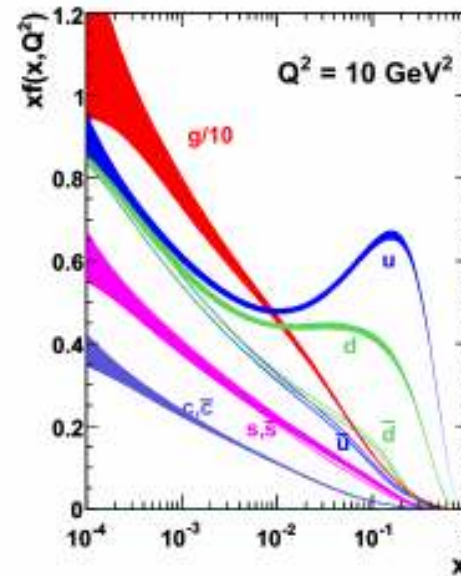
Higher twist terms

$$c_2(\rho) = c_2(0) + 4\pi^2 A_1^{q,s} M_N \rho - \frac{23}{9} \pi \alpha_s A_2^g M_N \rho$$

$$c_3(\rho) = c_3(0) - \frac{20}{3} \pi^2 A_3^{q,s} M_N^3 \rho$$

What has changed since 1992?

$$\begin{aligned}
 A_1^q &= 2 \int_0^1 dx x [q(x) + \bar{q}(x)] \\
 A_3^q &= 2 \int_0^1 dx x^3 [q(x) + \bar{q}(x)] \\
 A_1^s &= 2 \int_0^1 dx x [s(x) + \bar{s}(x)] \\
 A_3^s &= 2 \int_0^1 dx x^3 [s(x) + \bar{s}(x)] \\
 A_2^g &= \int_0^1 dx x g(x)
 \end{aligned}$$



A.D. Martin, W.J. Stirling, R.S. Thorne and G. Watt,
Eur. Phys. J. C **63**, 189 (2009).

1992

2013

$$A_1^q = 0.45$$

$$A_3^q = 0.06$$

$$A_1^s = 0.05$$

$$A_3^s = 0.002$$

$$A_2^g: \text{ignored}$$



$$A_1^q = 0.61$$

$$A_3^q = 0.065$$

$$A_1^s = 0.044$$

$$A_3^s = 0.0011$$

$$A_2^g = 0.359$$



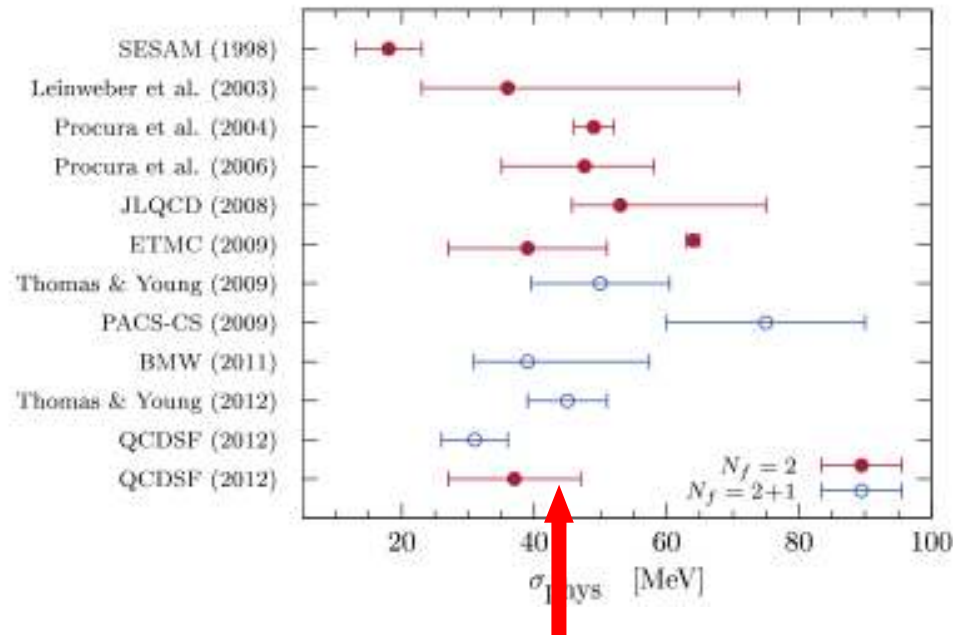
Some changes, but no
big effect.

non-negligible

What has changed since 1992?

The nuclear sigma term: $\sigma_{\pi N}$

Taken from G.S. Bali et al., Nucl. Phys. B866, 1 (2013).



Value used by Hatsuda and Lee: 45 MeV

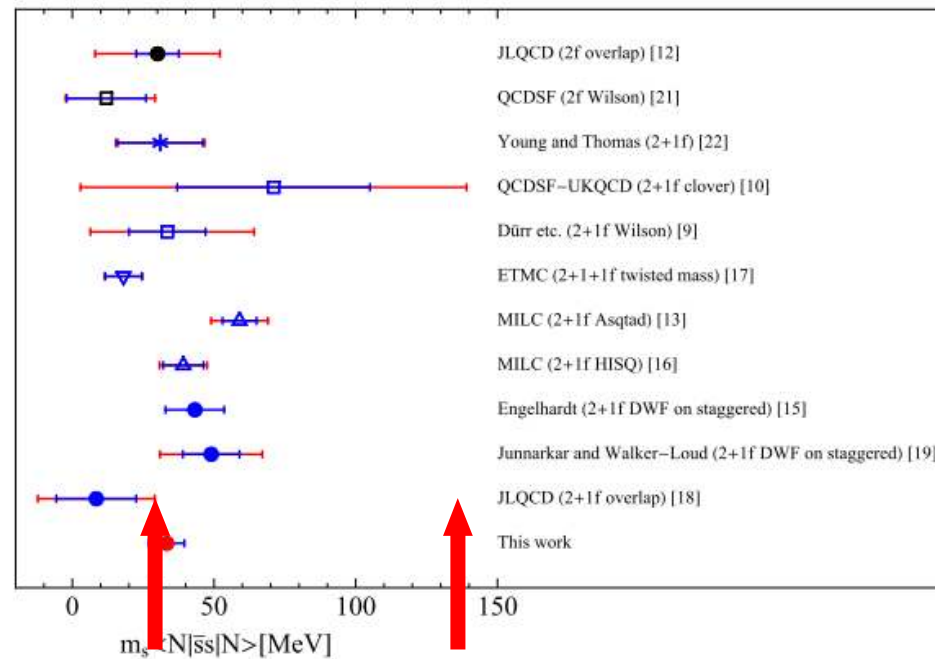
Recent lattice results are mostly consistent with the old values.
The latest trend might point to a somewhat smaller value.

What has changed since 1992?

The strangeness content of the nucleon:

$$y = \frac{\langle N | \bar{s}s | N \rangle}{\langle N | \bar{q}q | N \rangle}$$

Taken from M. Gong et al. (χ QCD Collaboration), arXiv:1304.1194 [hep-ph].



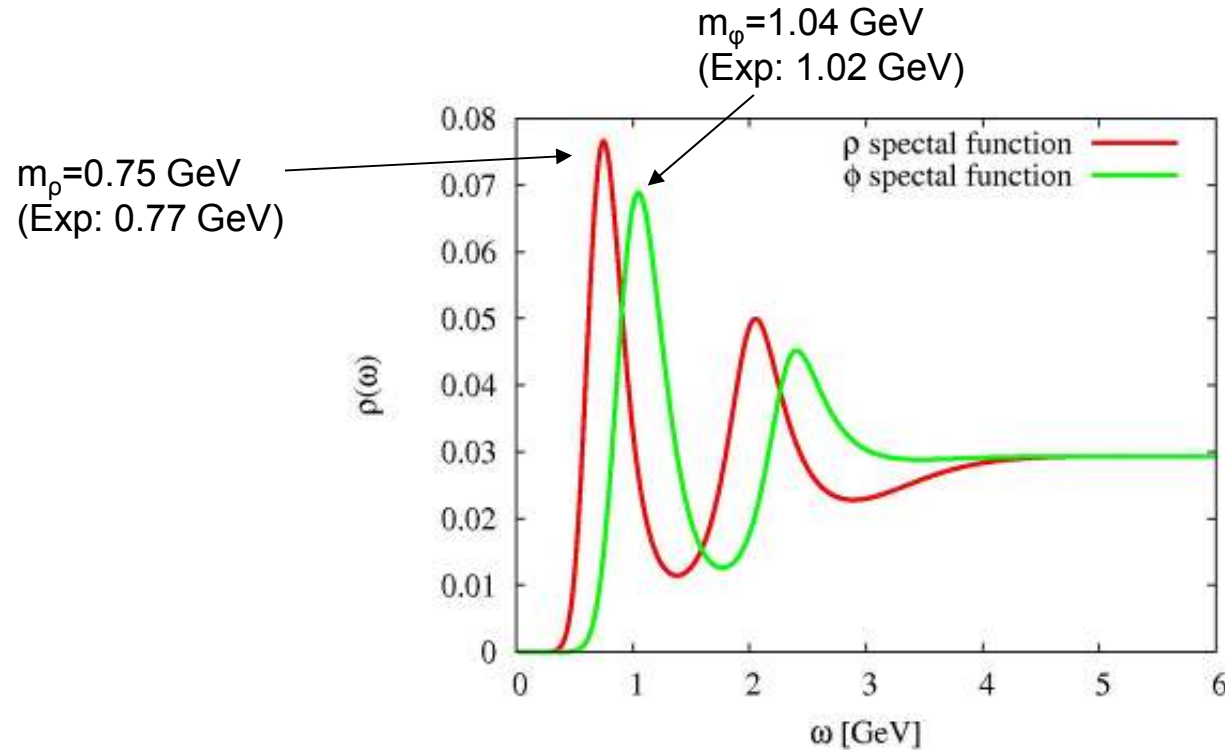
$y \sim 0.04$

Value used by Hatsuda and Lee: $y=0.22$
Too big!!

The value of y has shrunk by a factor of about 5: a new analysis is necessary!

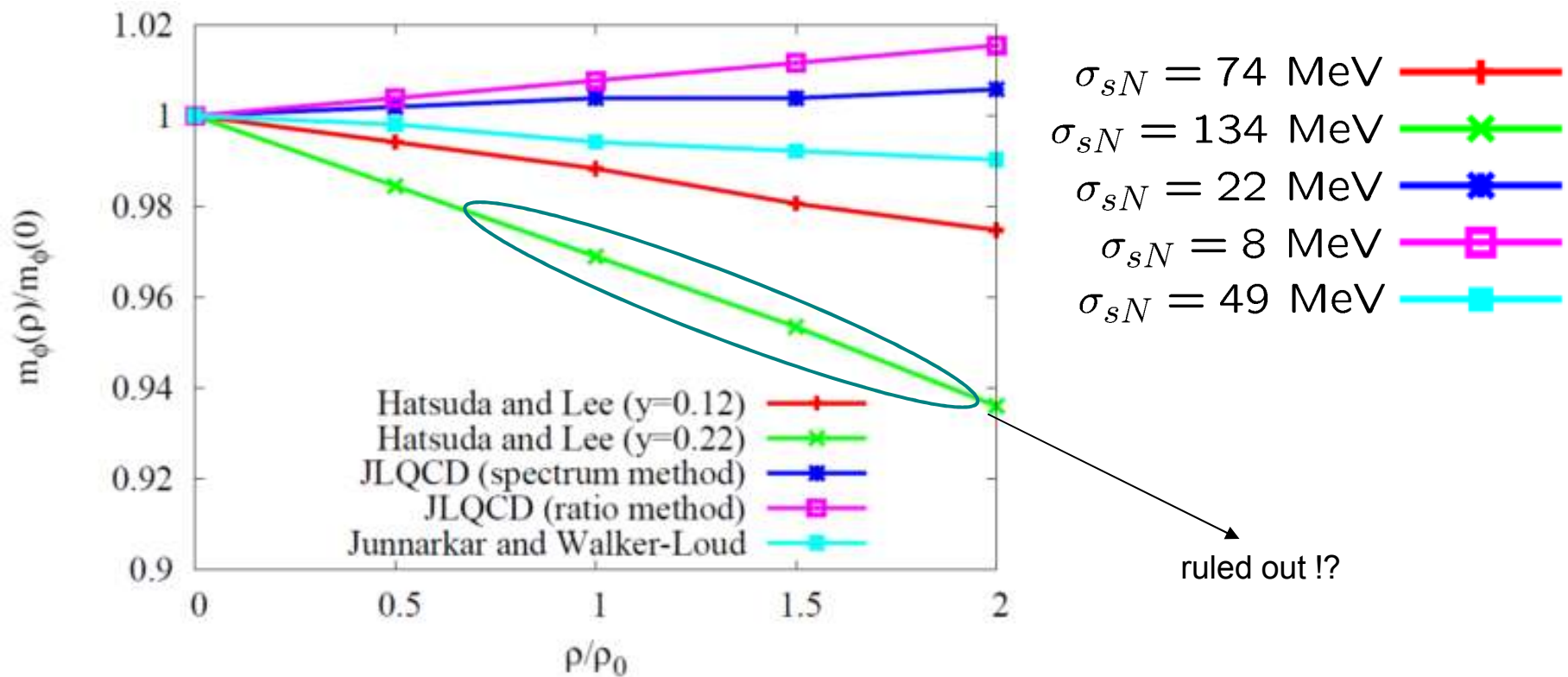
Results for the vacuum case

Analysis of the sum rule is done using the maximum entropy method (MEM), which allows to extract the spectral function from the sum rules without any phenomenological ansatz.



Used input parameters: $m_q = 4.7^{+0.9}_{-0.3} [\text{MeV}]$ $\langle \bar{q}q \rangle_0 = -(0.232 \pm 0.06)^3 [\text{GeV}^3]$
 $m_s = 128 \pm 7 [\text{MeV}]$ $\langle \bar{s}s \rangle_0 = (0.6 \pm 0.1) \langle \bar{q}q \rangle_0$
 $\alpha_s = 0.5$ $\langle \frac{\alpha_s}{\pi} G^2 \rangle_0 = 0.012 \pm 0.0036 [\text{GeV}^4]$

ϕ meson at finite density



The ϕ meson mass shift strongly depends on the strange sigma term.

What happened?

Let us examine the OPE at finite density more closely:

$$c_2(\rho) =$$

$$c_2(0) + \rho \left[-\frac{2}{27} M_N^0 + 2m_s \langle N | \bar{s}s | N \rangle + A_1^s M_N - \frac{23}{36} \frac{\alpha_s}{\pi} A_2^g M_N \right]$$


-65 MeV



16 MeV
98 MeV

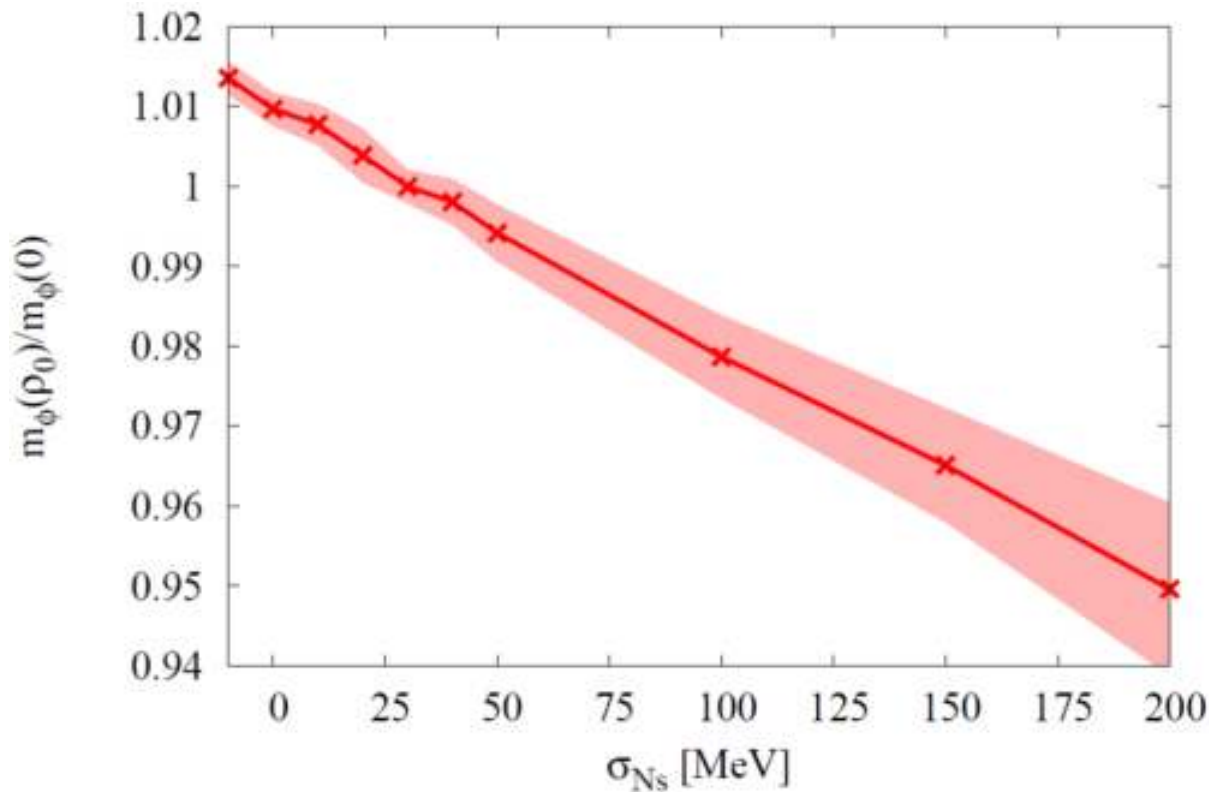

41 MeV


-34 MeV



Measuring the ϕ meson mass shift in nuclear matter provides a strong constraint to the strange sigma term!

Relation between the φ meson mass shift and the strange sigma term



$$\frac{m_\phi(\rho)}{m_\phi(0)} - 1 =$$

$$\left[b_0 + b_1 \left(\frac{\sigma_{sN}}{1 \text{ MeV}} \right) \right] \frac{\rho}{\rho_0}$$

$$b_0 = (1.24 \pm 0.36) 10^{-2}$$

$$b_1 = (-3.23 \pm 0.49) 10^{-4}$$

$$\frac{m_\phi(\rho)}{m_\phi(0)} - 1 =$$

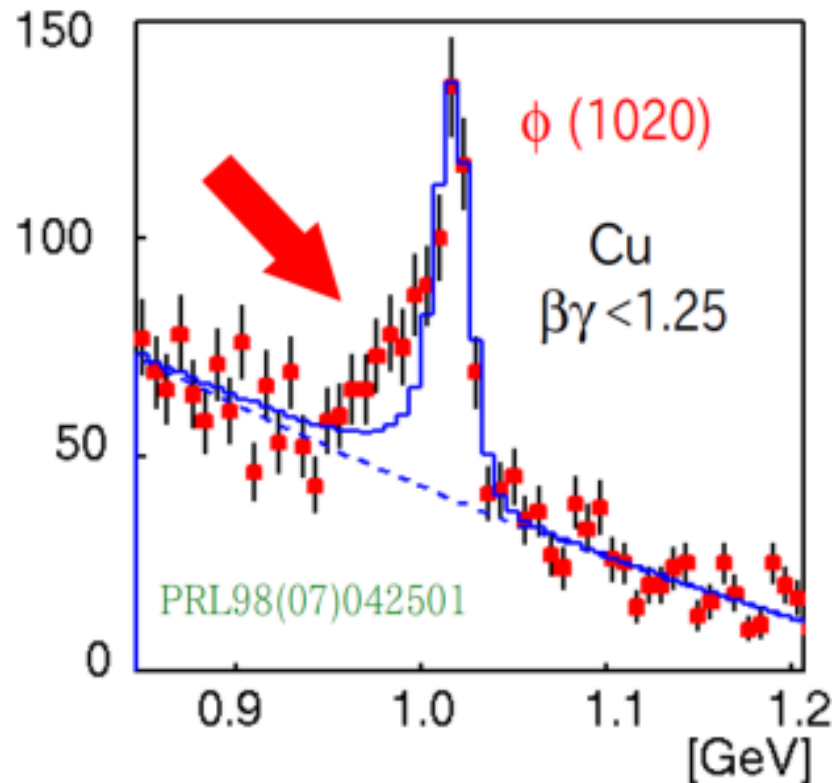
$$\left[b_0 + c_1 y \right] \frac{\rho}{\rho_0}$$

$$c_1 = (-1.97 \pm 0.30) 10^{-1}$$

However...

Experiments seem to suggest something else:

Result of the E325 experiment at KEK



35 MeV mass reduction
of the ϕ meson at nuclear
matter density!

R. Muto et al, Phys. Rev. Lett. 98, 042501 (2007).

What could be wrong?

1. So far neglected condensates

Terms containing higher orders of m_s and other so far neglected terms could have a non-negligible effect.

$$m_s^3 \langle \bar{s}s \rangle, \quad m_s^2 \langle \frac{\alpha_s}{\pi} G^2 \rangle, \quad m_s \langle \bar{s} g \sigma G s \rangle, \quad \dots$$

These terms do not have significant effects

2. α_s corrections

These corrections seem to be small

3. Underestimated density dependence of four-quark condensates

$$\langle \alpha_s (\bar{s} \gamma_\mu \gamma_5 \lambda^a s)^2 \rangle + \frac{2}{9} \langle \alpha_s (\bar{s} \gamma_\mu \lambda^a s) \sum_{q=u,d,s} (\bar{q} \gamma^\mu \lambda^a q) \rangle$$

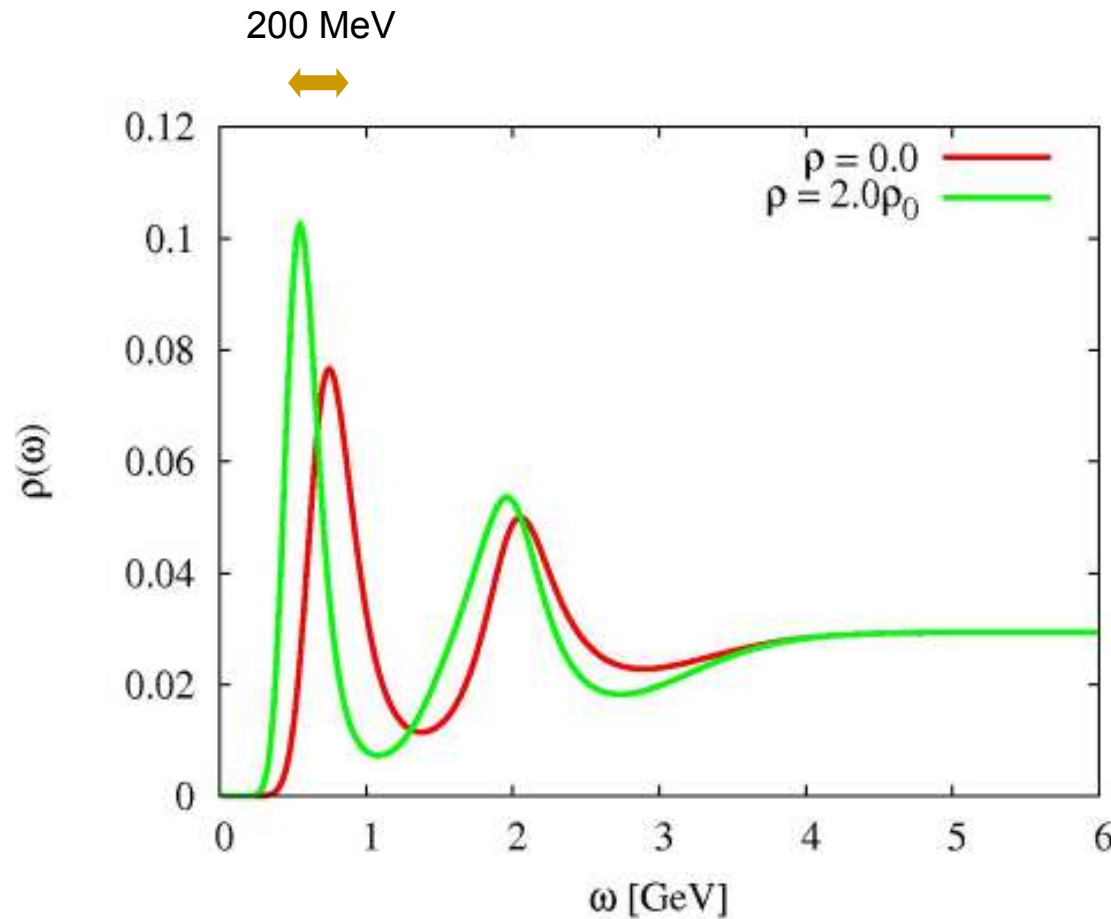
This should be checked in the future

Conclusions

- We have reanalyzed the light vector meson sum rules at finite density using MEM and the newest sigma-term values
 - For the ϕ -meson, due to the small strangeness content of the nucleon, the mass shift might be smaller than previously thought
 - The ϕ -meson mass shift at finite density is found to be strongly correlated with the strangeness content of the nucleon
 - Further study on the reliability of the obtained results are in progress
-

Backup slide

First results (ρ meson at finite density)



$$\frac{m_\rho(\rho)}{m_\rho(0)} \simeq 1 - C \frac{\rho}{\rho_0}$$

0.12 ~ 0.19

Consistent with the
result of Hatsuda-Lee.

Used input parameters: $\sigma_{\pi N} = 32 \pm 2 [\text{MeV}]$
 $M_N^0 = 880 [\text{MeV}]$

A. Semke and M.F.M. Lutz, Phys. Lett. B **717**, 242 (2012).

M. Procura, T.R. Hemmert and W. Weise,
 Phys. Rev. D **69**, 034505 (2004).

Estimation of the error of G(M)

$$G_{OPE}(M) = \frac{1}{4\pi^2} \left(1 + \frac{\alpha_s}{\pi} \right) + \left(2m \langle \bar{q}q \rangle + \frac{1}{12} \langle \frac{\alpha_s}{\pi} G^2 \rangle \right) \frac{1}{M^4} - \frac{112\pi}{81} \alpha_s \kappa \langle \bar{q}q \rangle^2 \frac{1}{M^6}$$

Gaussianly distributed values for the various parameters are randomly generated. The error is extracted from the resulting distribution of $G_{OPE}(M)$.

D.B. Leinweber, Annals Phys. **322**, 1949 (1996).

